Hard Scattering in Hadron-Hadron Collisions: Physics and Anatomy

#### **Section 8: Data Analysis Challenges**

- **1. Some Tools to Extract Knowledge**
- 2. Systematic Uncertainties
- 3. Significance (or not)
- 4. Perils of Running Blind

### **1. Introduction: Some Tools**

- Our understanding of high energy hadron collisions has limits
  - It's why we are studying them in the first place
  - But some of the limitations in knowledge "get in the way"
  - Progress is made by being able to control or minimize the uncertainties that issues not relevant to your analysis
- Generally, particle physicists have become pretty good at doing basic statistics
  - But we do get into trouble
  - Discuss a number of tools (and pitfalls) in common use

- Treatment of systematic uncertainties
  - Essential, but often riddled with assumptions and approximations
- Significance how do we make statements about belief from data?
  - But we do get into trouble
- Blind Analyses
  - All about avoiding unconscious or conscious bias
  - But there are challenges
- Resources Available
  - No re-invention of wheels please

## **Literature Summary**

#### Some classic statistics resources

- F. Solmitz, "Analysis of Experiments in Particle Physics", Annu. Rev. Nucl. Sci. 1964:14, 375-402.
- J. Orear, "Notes on Statistics for Physicists", CLNS 82/511 (1982), <u>http://pages.physics.cornell.edu/p510/w/images/p510b/6/62/</u> <u>Notes\_on\_Statistics\_for\_Physicists.pdf</u>

#### Systematic Uncertainty References

 P. Sinervo, "Definition and Treatment of Systematic Uncertainties", http://www.slac.stanford.edu/econf/C030908/papers/TUAT004.pdf

### 2. Systematic Uncertainties

- Systematic uncertainties play key role in physics measurements
  - Few formal definitions exist, much "oral tradition"
  - "Know" they are different from statistical uncertainties

#### **Random Uncertainties**

- Arise from stochastic fluctuations
- Uncorrelated with previous measurements
- □ Well-developed theory
- **Examples** 
  - measurement resolution
  - **inite statistics**
  - random variations in system

#### **Systematic Uncertainties**

- Due to uncertainties in the apparatus or model
- Usually correlated with previous measurements
- □ Limited theoretical framework
- □ Examples
  - **calibrations uncertainties**
  - detector acceptance
  - poorly-known theoretical parameters

## **Literature Summary**

#### Increasing literature on the topic of "systematics" A representative list:

- R.D.Cousins & V.L. Highland, NIM A320, 331 (1992).
- C. Guinti, Phys. Rev. D **59** (1999), 113009.
- G. Feldman, "Multiple measurements and parameters in the unified approach," presented at the FNAL workshop on Confidence Limits (Mar 2000).
- R. J. Barlow, "Systematic Errors, Fact and Fiction," hep-ex/0207026 (Jun 2002), and several other presentations in the Durham conference.
- G. Zech, "Frequentist and Bayesian Confidence Limits," Eur. Phys. J, C4:12 (2002).
- R. J. Barlow, "Asymmetric Systematic Errors," hep-ph/0306138 (June 2003).
- A. G. Kim et al., "Effects of Systematic Uncertainties on the Determination of Cosmological Parameters," astro-ph/0304509 (April 2003).
- J. Conrad et al., "Including Systematic Uncertainties in Confidence Interval Construction for Poisson Statistics," Phys. Rev. D 67 (2003), 012002
- G.C.Hill, "Comment on "Including Systematic Uncertainties in Confidence Interval Construction for Poisson Statistics"," Phys. Rev. D 67 (2003), 118101.
- G. Punzi, "Including Systematic Uncertainties in Confidence Limits", CDF Note in preparation.

#### **Case Study #1: W Boson Cross Section**

### Rate of W boson production

- Count candidates  $N_s + N_b$
- Estimate background
  - $N_b$  & signal efficiency e

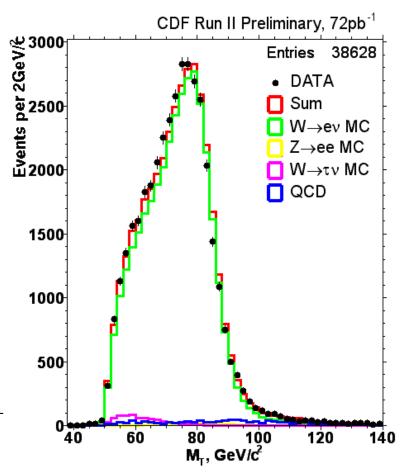
 $\sigma = \left(N_c - N_b\right) / (\varepsilon L)$ 

Measurement reported as

$$\sigma = 2.64 \pm 0.01 \text{ (stat)} \\ \pm 0.18 \text{ (syst) nb}$$

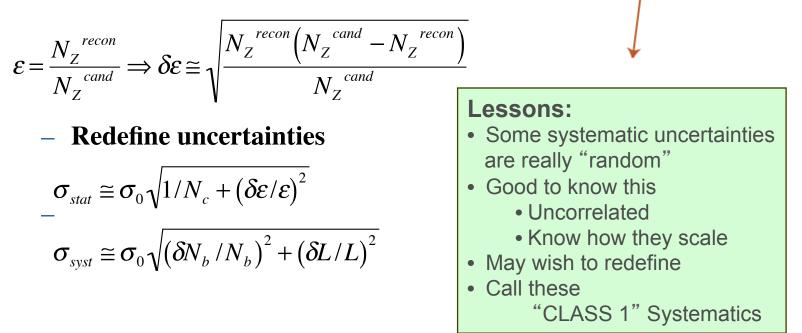
– Uncertainties are

$$\sigma_{stat} \cong \sigma_0^{stat} \sqrt{1/N_c}$$
$$\sigma_{syst} \cong \sigma_0^{syst} \sqrt{\left(\delta N_b / N_b\right)^2 + \left(\delta \varepsilon / \varepsilon\right)^2 + \left(\delta L / L\right)^2}$$



### **Definitions are Relative**

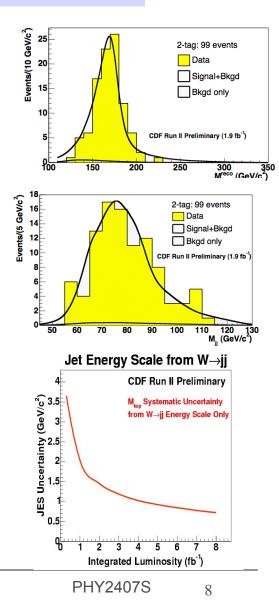
- Efficiency uncertainty estimated using Z boson decays
  - Count up number of Z candidates  $N_Z^{cand}$ 
    - > Can identify using charged tracks
    - > Count up number reconstructed  $N_Z^{recon}$



## **Top Mass Good Example**

- Top mass uncertainty in template analysis
  - Statistical uncertainty from shape of reconstructed mass distribution and statistics of sample
  - Systematic uncertainty coming from jet energy scale (JES)
    - Determined by calibration studies, dominated by modelling uncertainties
    - > 5% systematic uncertainty
- Latest techniques determine JES uncertainty from dijet mass peak (W->jj)
  - Turn JES uncertainty into a largely statistical one
  - Introduce other smaller systematics

 $M_{top} = 171.8 \pm 1.9 (\text{stat} + \text{JES}) \pm 1.0 (\text{syst}) \text{ GeV/c}^2$ = 171.9 ± 2.1 GeV/c<sup>2</sup>

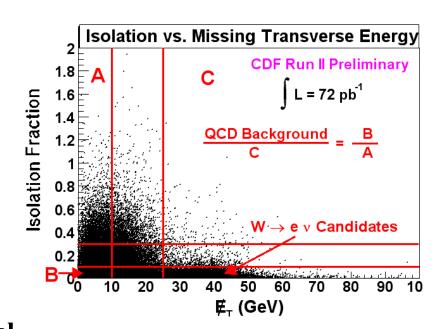


#### **Case Study #2: Background Uncertainty**

#### Look at same W cross section analysis

#### - Estimate of $N_b$ dominated by QCD backgrounds

- > Candidate event
  - Have non-isolated leptons
  - Less missing energy
- > Assume that isolation and MET uncorrelated
- > Have to estimate the uncertainty on  $N_b^{QCD}$
- No direct measurement
   has been made to verify the model
- Estimates using Monte Carlo modelling have large uncertainties



## **Estimation of Uncertainty**

### Fundamentally different class of uncertainty

- Assumed a model for data interpretation
- Uncertainty in  $N_b^{QCD}$  depends on accuracy of model
- Use "informed judgment" to place bounds on one's ignorance
  - > Vary the model assumption to estimate robustness
  - > Compare with other methods of estimation

### Difficult to quantify in consistent manner

– Largest possible variation?

> Asymmetric?

– Estimate a "1  $\sigma$ " interval?

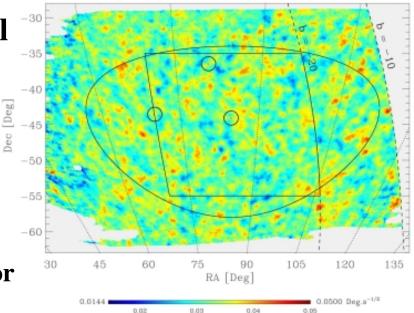
**Take** 
$$\sigma \approx \frac{\Delta}{\sqrt{12}}$$
?

#### Lessons:

- Some systematic uncertainties reflect ignorance of one's data
- Cannot be constrained by observations
- Call these
   "CLASS 2" Systematics

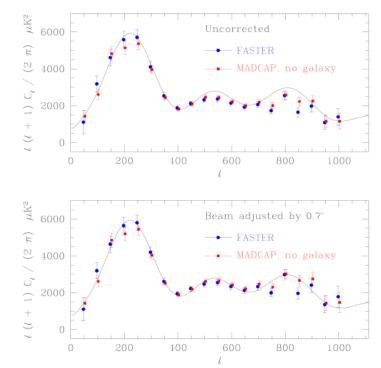
#### Case Study #3: Boomerang CMB Analysis

- Boomerang is one of several CMB probes
  - Mapped CMB anisoptropy
  - Data constrain models of the early universe
- Analysis chain:
  - Produce a power spectrum for the CMB spatial anisotropy
    - > Remove instrumental effects through a complex signal processing algorithm
  - Interpret data in context of many models with unknown parameters



#### **Incorporation of Model Uncertainties**

- Power spectrum extraction includes all instrumental effects
  - Effective size of beam
  - Variations in data-taking procedures
- Use these data to extract7 cosmological parameters
  - Take Bayesian approach
    - > Family of theoretical models defined by 7 parameters
    - > Define a 6-D grid (6.4M points), and calculate likelihood function for each

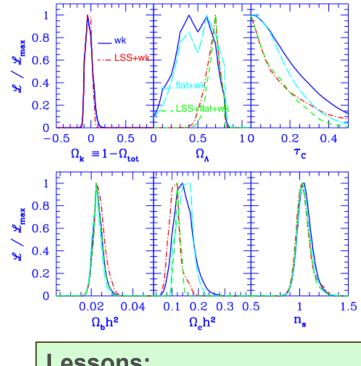


### **Marginalize Posterior Probabilities**

- **Perform a Bayesian** "averaging" over a grid of parameter values
  - Marginalize w.r.t. the other parameters
    - > NB: instrumental uncertainies included in approximate manner
  - Chose various priors in the parameters

### **Comments:**

- Purely Bayesian analysis with no frequentist analogue
- **Provides path for inclusion of** additional data (eg. WMAP)



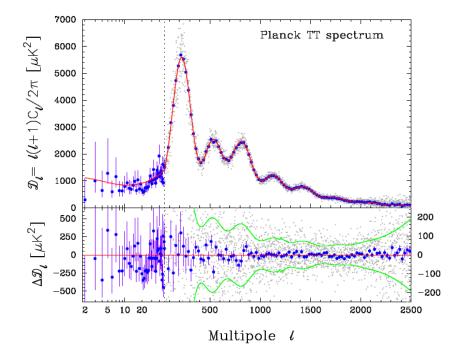
#### Lessons:

- Some systematic uncertainties reflect paradigm uncertainties
- · No relevant concept of a frequentist ensemble
- Call these "CLASS 3" Systematics

### **Latest Planck Results**

#### The prior uncertainties dominate

		Planck			
Parameter	Best fit	68% limits			
$\Omega_{ m b}h^2$	0.022068	$0.02207 \pm 0.00033$			
$\Omega_{\rm c} h^2$	0.12029	$0.1196 \pm 0.0031$			
100 <i>θ</i> <sub>мс</sub>	1.04122	$1.04132 \pm 0.00068$			
τ	0.0925	$0.097 \pm 0.038$			
<i>n</i> <sub>s</sub>	0.9624	$0.9616 \pm 0.0094$			
$\ln(10^{10}A_{\rm s})$	3.098	$3.103\pm0.072$			
$\Omega_{\Lambda}$	0.6825	$0.686 \pm 0.020$			
$\Omega_m$	0.3175	$0.314 \pm 0.020$			
<i>σ</i> <sub>8</sub>	0.8344	$0.834 \pm 0.027$			
Z <sub>re</sub>	11.35	$11.4^{+4.0}_{-2.8}$			
$H_0$	67.11	$67.4 \pm 1.4$			
10 <sup>9</sup> A <sub>s</sub>	2.215	$2.23\pm0.16$			
$\Omega_{\rm m} h^2 \dots$	0.14300	$0.1423 \pm 0.0029$			
$\Omega_{\rm m}h^3$	0.09597	$0.09590 \pm 0.00059$			
<i>Y</i> <sub>P</sub>	0.247710	$0.24771 \pm 0.00014$			
Age/Gyr	13.819	$13.813\pm0.058$			



Planck Collaboration, 1303.5076v3 (2014)

### Proposed Taxonomy for Systematic Uncertainties

- Three "classes" of systematic uncertainties
  - Uncertainties that can be constrained by ancillary measurements
  - Uncertainties arising from model assumptions or problems with the data that are poorly understood
  - Uncertainties in the underlying models
- Estimation of Class 1 uncertainties straightforward
  - Class 2 and 3 uncertainties present unique challenges
  - In many cases, have nothing to do with statistical uncertainties
    - > Driven by our desire to make inferences from the data using specific models

## **Estimation Techniques**

- No formal guidance on how to define a systematic uncertainty
  - Can identify a possible source of uncertainty
  - Many different approaches to estimate their magnitude
    - > Determine maximum effect D
- General rule:
  - Maintain consistency with definition of statistical intervals

$$\sigma = \frac{\Delta}{\sqrt{12}}?$$

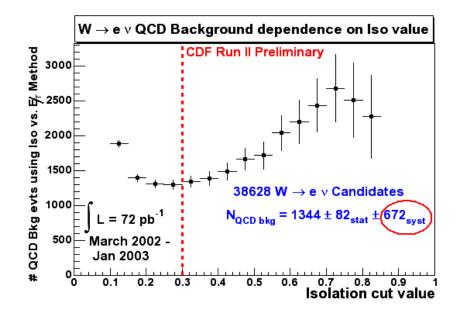
 $\sigma = \frac{\Delta}{2}?$ 

- Field is pretty glued to 68% confidence intervals
- Recommend attempting to reflect that in magnitudes of systematic uncertainties
- Avoid tendency to be "conservative"

#### Estimate of Background Uncertainty in Case Study #2

### Look at correlation of Isolation and MET

- Background estimate increases as isolation "cut" is raised
- Difficult to measure or accurately model
  - Background comes primarily from very rare jet events with unusual properties
  - > Very model-dependent



#### Assume a systematic uncertainty representing the observed variation

- Authors argue this is a "conservative" choice

### **Cross-Checks Vs Systematics**

### **R.** Barlow makes the point in Durham(PhysStat02)

- A cross-check for robustness is not an invitation to introduce a systematic uncertainty
  - > Most cross-checks confirm that interval or limit is robust,
    - They are usually not designed to measure a systematic uncertainty
- More generally, a systematic uncertainty should
  - Be based on a hypothesis or model with clearly stated assumptions
  - Be estimated using a well-defined methodology
  - Be introduced *a posteriori* only when all else has failed

### **Statistics of Systematic Uncertainties**

- Goal has been to incorporate systematic uncertainties into measurements in coherent manner
  - Increasing awareness of need for consistent practice
    - > Frequentists: interval estimation increasingly sophisticated
      - Neyman construction, ordering strategies, coverage properties
    - > Bayesians: understanding of priors and use of posteriors
      - Objective vs subjective approaches, marginalization/conditioning
  - Systematic uncertainties threaten to dominate as precision and sensitivity of experiments increase
  - There are a number of approaches widely used
    - Summarize and give a few examples
    - Place it in context of traditional statistical concepts

### **Formal Statement of the Problem**

Have a set of observations x<sub>i</sub>, i=1,n

- Associated probability distribution function (pdf) and likelihood function  $p(x_i | \theta) \Rightarrow \mathcal{L}(\theta) = \prod_i p(x_i | \theta)$ 

> Depends on unknown random parameter *q* 

> Have some additional uncertainty in pdf

– Introduce a second unknown parameter /

 $\mathcal{L}(\theta,\lambda) = \prod_{i} p(x_i \mid \theta,\lambda)$ 

In some cases, one can identify statistic y<sub>j</sub> that provides information about /

$$\mathcal{L}(\boldsymbol{\theta}, \boldsymbol{\lambda}) = \prod_{i, j} \mathbf{p}(x_i, y_j \mid \boldsymbol{\theta}, \boldsymbol{\lambda})$$

- Can treat / as a "nuisance parameter"

## **Bayesian Approach**

Identify a prior p(l) for the "nuisance parameter" /

- Typically, parametrize as either a Gaussian pdf or a flat distribution within a range ("tophat")
- Can then define Bayesian posterior

 $\mathcal{L}(\theta,\lambda) \, \pi(\lambda) \, d\theta \, d\lambda$ 

- Can marginalize over possible values of /
  - > Use marginalized posterior to set Bayesian credibility intervals, estimate parameters, etc.

### Theoretically straightforward ....

- Issues come down to choice of priors for both q, /
  - > No widely-adopted single choice
  - > Results have to be reported and compared carefully to ensure consistent treatment

### **Frequentist Approach**

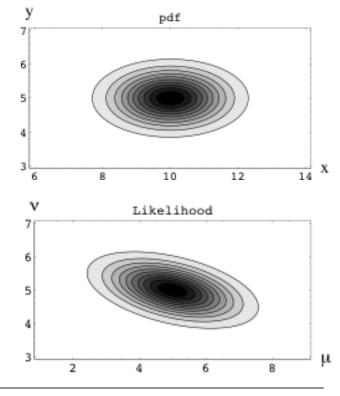
- **Start with a pdf for data**  $p(x_i, y_j | \theta, \lambda)$ 
  - In principle, this would describe frequency distributions of data in multi-dimensional space
  - Challenge is take account of nuisance parameter
  - Consider a toy model

$$p(x, y \mid \mu, v) = G(x - (\mu + v), 1)G(y - v, s)$$

Parameter s is Gaussian
 width for n

### ■ Likelihood function (*x=10*, *y=5*)

- Shows the correlation
- Effect of unknown n



#### Formal Methods to Eliminate Nuisance Parameters

- Number of formal methods exist to eliminate nuisance parameters
  - Of limited applicability given the restrictions
  - Our "toy example" is one such case
    - > Replace *x* with *t*=*x*-*y* and parameter *n* with

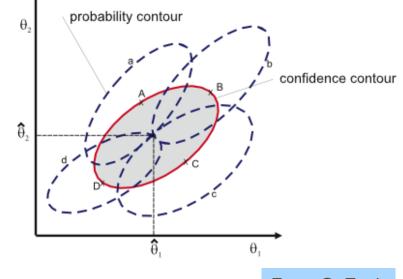
$$v' \equiv v + \frac{\mu s^2}{1 + s^2}$$
  
$$\Rightarrow p(t, y \mid \mu, v') = G(t - \mu, \sqrt{1 + s^2})G(y - v' + \frac{ts^2}{1 + s^2}, \frac{s}{\sqrt{1 + s^2}})$$

> Factorized pdf and can now integrate over *n*'

- > Note that pdf for *m* has larger width, as expected
- In practice, one often loses information using this technique

#### Alternative Techniques for Treating Nuisance Parameters

- Project Neyman volumes onto parameter of interest
  - "Conservative interval"
  - Typically over-covers, possibly badly
- Choose best estimate of nuisance parameter
  - Known as "profile method"
  - Coverage properties require definition of ensemble



From G. Zech

- Can possible under-cover when parameters strongly correlated
  - Feldman-Cousins intervals tend to over-cover slightly (private communication)

#### **Example: Solar Neutrino Global Analysis**

#### Many experiments have measured solar neutrino flux

- Gallex, SuperKamiokande, SNO, Homestake, SAGE, etc.
- Standard Solar Model (SSM) describes n spectrum
- Numerous "global analyses" that synthesize these

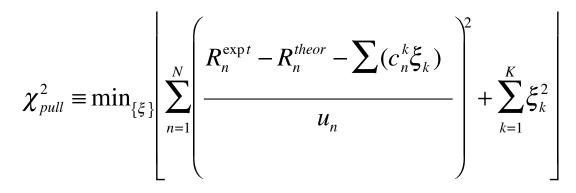
#### **Fogli et al. have detailed one such analysis**

- 81 observables from these experiments
- Characterize systematic uncertainties through 31 parameters
  - > 12 describing SSM spectrum
  - > 11 (SK) and 7 (SNO) systematic uncertainties
- **Perform a**  $\chi^2$  analysis
  - Look at  $\chi^2$  to set limits on parameters

Hep-ph/0206162, 18 Jun 2002

## **Formulation of** $\chi^2$

# In formulating $\chi^2$ , linearize effects of the systematic uncertainties on data and theory comparison



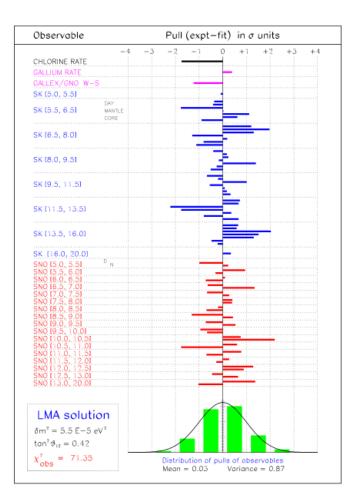
> Uncertainties  $u_n$  for each observable

#### - Introduce "random" pull $x_k$ for each systematic

- > Coefficients  $c_k^n$  to parameterize effect on *nth* observable
- > Minimize  $\chi^2$  with respect to  $x_k$
- > Look at contours of equal  $\Delta \chi^2$

## **Solar Neutrino Results**

- Can look at "pulls" at  $\chi^2$  minimum
  - Have reasonable distribution
  - Demonstrates consistency of model with the various measurements
  - Can also separate
    - > Agreement with experiments
    - > Agreement with systematic uncertainties



### **Pull Distributions for Systematics**

### Pull distributions for x<sub>k</sub> also informative

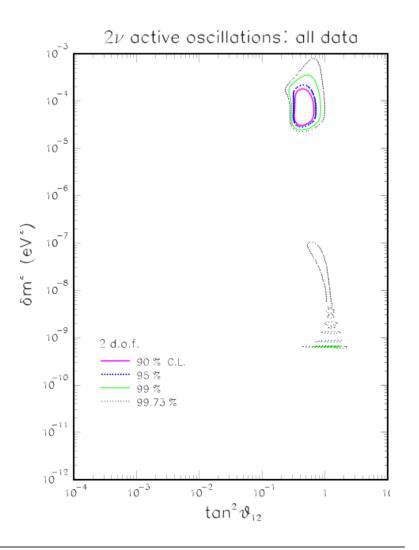
- Unreasonably small variations
- Estimates are globally too conservative?
- Choice of central values affected by data
  - Note this is NOT a blind analysis
- But it gives us some confidence that intervals are realistic

Systematics	Pulls	(σ)	for	LMA	solu	itio	n
		-4	-3 -	2 -1	0 +1	+2	+3 +4
S11	-0.05						
S33	+0						
S34	+0.01						
S1,14	-0.15				•		
S17	+0.38				•		
Luminosity	+0.04						
Z/X	+0.03						
Age	+0						
Opacity	-0.05						
Diffusion	-0.02						
CBe	-0.07						
Shep	-0.03						
8B shape	+0.17				•		
SK scale	+0.78				-		
SK resol.	+0.61				-		
SK offset	+0.44				-		
SK [5.0, 5.5]	-0.03						
SK [5.5, 6.5]	-0.26				•		
SK [6.5, 8.0]	+0.54				-		
SK [8.0, 9.5]	+0.01						
SK [9.5, 11.5]	-0.14				•		
SK [11.5, 13.5]	-0.21				•		
SK [13.5, 16.0]	+0.26				•		
SK [16.0, 20.0]	+0.01						
SNO scale	-0.15				•		
SNO resol.	-0.32				•		
SNO vertex	+0.13				•		
SNO n capture	-0.1						
SNO n bkgd	-0.06						
SNO LE bkgd	-0.16				•		
SNO cross sec.	+0.04						
				$\chi^2_{\rm sys}$	= 2.0	5	

## **Typical Solar Neutrino Contours**

#### Can look at probability contours

- Assume standard  $\chi^2$  form
- Probably very small probability contours have relatively large uncertainties



## **Hybrid Techniques**

- A popular technique (Cousins-Highland) does an "averaging" of the pdf
  - Assume a pdf for nuisance parameter g(*l*)
  - "Average" the pdf for data x

 $p_{CH}(x \mid \theta) \equiv \int p(x \mid \theta, \lambda) g(\lambda) d\lambda$ 

- Argue this approximates an ensemble where
  - > Each measurement uses an apparatus that differs in parameter /
    - The pdf g(*l*) describes the frequency distribution
  - > Resulting distribution for x reflects variations in /

### Intuitively appealing

- But fundamentally a Bayesian approach
- Coverage is not well-defined

See, for example, J. Conrad et al.

## **Computationally Challenging**

#### In many measurements

- Can have several dozen sources of systematic uncertainty
- Creating a tractable ensemble is not possible
- Even the definition of the ensemble is controversial
- Current state of the art is to perform a Bayesian-like "marginalization"
  - Treat the new probability function in the same way as before
  - But
    - > Not clear how to evaluate coverage
    - > Not strongly grounded in theory

## 3. What is Significance?

### Typical HEP approach

- Have a set of observations
- We say the data are "statistically significant" when
  - > We can use data to support a specific hypothesis, eg.
    - "We see a phenomenom not predicted by the Standard Model"
    - "We report the discovery of X"
  - > The interpretation eliminates a number of competing hypotheses
  - > The conclusion will not likely be altered with larger statistics or further analysis
- Want a statistical framework that
  - Measures "degree of belief"
  - Ensures robust conclusions

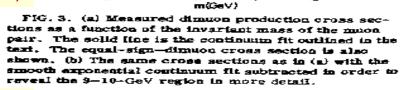
## **Some "Obvious" Discoveries**

#### **Observation of B<sup>o</sup>B**<sup>o</sup>Mixing - 24.8 ± 7.6 ± 3.8 like-sign events vs $25.2 \pm 5.0 \pm 3.8$ opposite sign drayly-0 (cm<sup>1</sup>/JeV/nuclean) - " $3\sigma$ " discovery Albrecht et al., PLB 192, 245 (1987) **W** Boson – 6 ev events, no background! Arnison et al., Upsilon PLB 122, 103 (1983) m(GeV) 770 events on 350 background ο, <sub>ino</sub>lio<sup>17</sup> an<sup>t</sup>Get/inclear - Described as "significant" but no measure of it Herb et al., PRL 39, 252 (1977) **B** mesons

#### 18 avanta an 47 h

- 18 events on 4-7 background
- No measure of significance

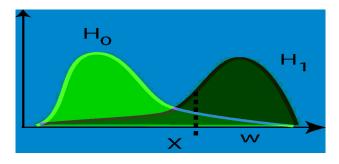
Behrends et al., PRL 50, 881 (1983)



## **A Frequentist Definition**

### Significance defined in context of "hypothesis testing"

- Have two hypotheses, H<sub>0</sub> and H<sub>1</sub>, and possible set of observations X
  - > Choose a "critical region", w, in the space of observations X
  - > Define **significance**,  $\alpha$ , as probability of X  $\in$  w when H<sub>0</sub> is true
  - > Define the **power**, 1- $\beta$ , as probability of X  $\in$  w when H<sub>1</sub> is true



Typically, H0 is "null" hypothesis

In this language, an observation is "significant" when

- Significance  $\alpha$  is small &  $\beta$  is small
  - > Typically a < few  $10^{-5}$

## **Some Comments on Formal Definition**

### Definition depends on

#### Choice of statistic X

- > Left up to the experimenter as part of design
- > More on that later

#### Choice of "critical region" w

- > Depends on hypotheses
- > Often chosen to minimize systematic uncertainties?
- > Not necessarily defined in advance!

#### Definition of "probability"

- > A frequentist definition
- > Raises issue of how systematic uncertainties are managed

#### – Choice of $\alpha$ and $\beta$

- > Matter of "taste" and precedent
- > A small  $\alpha$  is safe, but comes with less "discovery reach"

### More fundamentally:

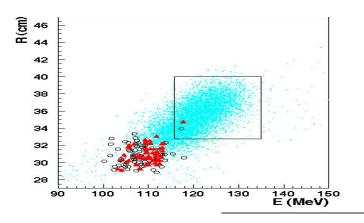
– Is this an adequate definition of "significance?"

## The Choice of Statistic & Critical Region

- Choice of statistic motivated by specific experimental design
  - Informed by the measurement to be made
  - Critical region is chosen at the same time
  - Good example: E787/E949 search

 $K^{\scriptscriptstyle +}\!\to\pi^{\scriptscriptstyle +}\nu\nu$ 

- > Look for  $\pi^+ \rightarrow \mu^+ \nu$  decay
- > Define a "box" a priori
  - Expected 0.15±0.05 event bkgd



Only two events Observed

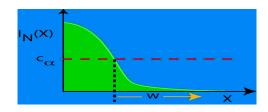
Significance 0.02%

Have used the "box" Since 1988

# **Optimal Tests: Neyman-Pearson**

- In some cases, possible to identify the "most powerful" test
  - Must involve only "simple" hypotheses (no free parameters)
    - > PDF's given by  $f_i(X)$
    - > Must have two hypotheses
  - For given  $\alpha$ , can identify region to minimize  $\beta$  for alternative  $H_1$ 
    - > Order observations by  $I_N(X) \equiv f_0(X) / f_1(X)$
    - > Can minimize  $\beta$  by choosing critical region as all X s.t.  $l_N(X) \ge c_a$

- Chose  $c_a$  so that  $\int_{M} \mathbf{f}_0(\mathbf{X}) d\mathbf{X} = \alpha$ 



# **Caveats to Neyman-Pearson**

### Neyman-Pearson limited

#### Only true for simple hypotheses

Not for composite hypotheses (where unknown parameter)

#### Compares two hypotheses

- > Depends on alternative hypothesis
- > Makes results model-dependent
- But does give some insight
  - $\ The \ ratio \ I_N(X) \ is \ proportional \ to \ ratio \ of \ likelihoods$

$$f_0(X) / f_1(X) \cong L_0(X) / L_1(X)$$

- Provides guidance for definition of effective tests

# **Definition of Critical Region**

# Challenge is not to bias choice of critical region with data

#### - However, observer required to understand data

- > Identify instrumental pathologies
- > Identify unexpected backgrounds
- > Estimate systematic uncertainties
- > Verify stable run conditions
- Studies may lead to unconscious bias (see, eg. RPP plots!)

### Blind" analyses are popular

- > Study data complementary to signal
- > However, implementation varies
  - SNO's pure  $D_2O$  results set aside about 40% of data
  - Not clear that this really helps!
- > Even E787/E949 reserve right to examine background rejection

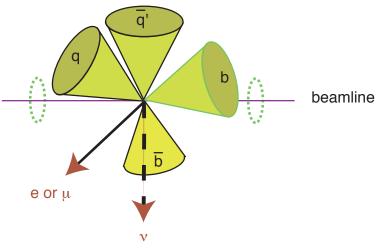
# Significance in Counting Experiments

### Top quark search is textbook example

- By 1991, CDF had ruled out top quark with mass < 91 GeV/c<sup>2</sup>
- Searching for top quark pair production and decay into
  - > Lepton + n + jets (20%)
  - > Dilepton + n + jets (8%)

### In a sample of 20 pb<sup>-1</sup>, expected handful of events

- Large background from W + jets
- "Fake" b-quark tags



### **Definition of the Measurement**

- Defined clear strategy in 1990
  - Identify lepton+jets and dilepton candidates
  - Count "b" tags in lepton+jet events
    - > Use two b-tagging algorithms
      - Use events with 1-2 jets as control
      - Signal sample events with ≥3 jets
      - Expected 3.5 evts (M<sub>top</sub>=160 GeV/c<sup>2</sup>)

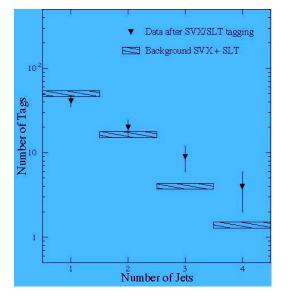
Expect **5.4±0.4** tags from background

Observed **13** tagged "b jets" in 10 evts

7 SVX tags 6 lepton tags

#### - For dileptons:

- > Require 2 or more jets
- > Expected 1.3 evts ( $M_{top}=160 \text{ GeV/c}^2$ )
- > Observed 2 evts, bkd of 0.6±0.3 evts



## **Significance Calculation**

#### Calculated probability of background hypothesis

- Dilepton significance  $\alpha_{dil} = 0.12$
- Used MC calculation
  - Treated background uncertainty as a normally distributed uncertainty on acceptance
- For lepton+jets, MC gives
  - > SVX b tags:  $\alpha_{SVX} = 0.032$
  - > SLT b tags:  $\alpha_{SLT} = 0.038$

#### To combine, take into account correlations

- Gives  $\alpha_{tot} = 0.0026$
- If assume independent, then

$$\alpha_{tot} = \alpha_{dil} \, \alpha_{ljets} \, \left[ 1 - \ln(\alpha_{dil} \, \alpha_{ljets}) \right]$$

> Gives  $\alpha_{tot} = 0.0088$ 

- Collaboration reported only "evidence for top quark...."
  - > Factor 2 more data --  $\alpha_{tot}$  = few 10<sup>-5</sup>

# **Power of the Top Quark Statistic**

- Choice of statistic driven by need to reduce background
  - Note  $\varepsilon_{ljets} = 0.074$  before b-tagging
    - > Predict 12 events signal and 60 events background
    - > Tagging efficiency 0.40
      - Background "efficiency" 0.09
  - Definition of "power" problematic
    - > Arbitrary
      - Power of lepton+jets selection? b-tagging?
      - A posteriori choice of  $X = N_{tags} + N_{dil}$
    - > Experimenter chooses "critical region" based on hypothesis
      - Lepton+jets Higgs search use $\delta$  different selection

 $WH \rightarrow l v b b$ 

- Usually characterized by sensitivity
  - > Size of expected signal

## Significance using Data Distributions

#### Measurements often involve continuous observables

- Can assess agreement with "null" hypothesis
  - > Generally "goodness-of-fit" tests

#### Number of tests in common use

>  $\chi^2$  Test

- Depends on choice of binning
- Limited to "large" statistics samples
  - Bin contents > 5-10 (?)
- > Smirnov-Cramer-Von Mises
  - Define statistic based on cumulative distributions  $S_N(x)$

$$W^{2} \equiv \int \left[ S_{N}(X) - F(X) \right]^{2} f(X) \, dX$$

- Probability distribution for W<sup>2</sup> independent of distribution
  - E[W<sup>2</sup>] = (6N)<sup>-1</sup> and V[W<sup>2</sup>] = (4N-3)/180N<sup>3</sup>

#### > Kolmogorov-Smirnov

- Popular form of test based on  $S_N(x)$
- Distribution for  $D_N$  proportional to  $\chi^2$

$$\mathsf{D}_{\mathsf{N}} \equiv \max \big| \mathsf{S}_{\mathsf{N}}(\mathsf{X}) - \mathsf{F}(\mathsf{X}) \big|$$

# **Multivariate Significance**

### Often difficult to reduce data to 1-dimensional statistic

#### - Typical case has several variables

- > Different correlations between signal and "null" hypothesis
- > Any straightforward transformation causes loss of information

#### Several techniques used

- > Characterize significance of each component and then combine into a single measure of significance
- > More sophisticated, e.g.
  - Combine information using any one of the techniques discussed by Prosper, Towers, etc.

#### In practice, two approaches:

- **1.** Assume independent statistics
  - Check for any correlations
- 2. Model correlations using MC approaches or "bootstrapping"
  - Computationally expensive
  - Relies on understanding correlations

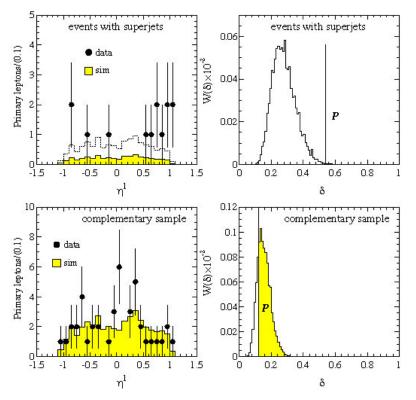
# An Infamous Example: "Superjets"

### **CDF Run I data contained**

- Unusual lepton + v + 2,3 jet events
  - > 13 events with jets that are both SLT and SVX tagged
    - Expect 4.4±0.6 events from background sources
    - Significance is 0.001!
- Led to examination of 9 kinematical distributions
  - $P_{T}$  &  $\eta$  for leptons & jets, and azimuthal angle between lepton, jet
  - $-~P_{T}$  and  $\eta$  for lepton+jet system
  - > Perform independent K-S tests
    - Use control sample defined by events without a "supertag"
    - Combined significance of 1.6x10<sup>-6</sup>
  - > Also defined a new statistic
    - Sum of K-S distances
    - MC gives significance of 3.3x10<sup>-6</sup>

# **K-S Tests on Superjet Data**

### Lepton η distribution



#### - Some approximations:

- > Control sample events w/o superjet
- > Randomly pick 13 of 42 events

# **Comments on Superjet Study**

### **Choice of statistic (number of superjets) problematic**

- Made *a posteriori* after anomaly noted
  - > Significance difficult to assess
- Ignored lepton + 1 jet data (where one observes a deficit of events)
   > Why?

#### Choice of distributions also problematic

- Justified *a posteriori*
- Correlations difficult to assess

#### Aside:

- Interpretation of excess requires unusual physics process
  - > Not a problem in itself
  - > But small statistics allow for many hypotheses

# Some Practical Proxies for Significance

### **HEP suffers Gaussian tyranny**

- Many people will quote numbers of " $\sigma$ " as measures of significance
  - > Belief that this can be more readily interpreted by lay person
    - Shorthand for the significance of an ns measurement
  - >  $5\sigma$  seems to have become conventional "discovery threshold"
    - $\alpha = 2.8 \times 10^{-7}$
    - Used for LHC discovery reach

### In situations where expected signal S and background B

#### - Various figures of merit

- > S/N -- signal versus noise
  - Doesn't scale with N

Just normal Gaussian estimate of # of s.d.

> More natural definition is

– Does scale with N

$$\sqrt{\mathsf{B}}$$

S

See papers by Bityukov & Krasnikov for more discussion

# The "Flip-Flopping" Physicist

- Feldman & Cousins highlighted the problem of "flip-flopping"
  - A physicist who uses
    - > One set of criteria to set a limit in the absence of a signal
    - > Different criteria to claim a significant signal
  - Results in confidence intervals with ill-defined frequentist coverage
- This should be anticipated in any experiment that wishes to be sensitive to small signals
  - F-C propose their "unified approach"

## What About Reverend Bayes?

Bayesian approach to classifying hypotheses is

$$\frac{\mathsf{P}(\mathsf{H}_1 \mid \mathsf{X})}{\mathsf{P}(\mathsf{H}_0 \mid \mathsf{X})} = \frac{\mathsf{P}(\mathsf{X} \mid \mathsf{H}_1)}{\mathsf{P}(\mathsf{X} \mid \mathsf{H}_0)} \bullet \frac{\pi(\mathsf{H}_1)}{\pi(\mathsf{H}_0)}$$

#### - Few comments:

- > P(XlH<sub>i</sub>) is typically likelihood
- > Only meaningful in comparison of two hypotheses
- > Can handle composite hypotheses readily
  - Just integrate over any "nuisance" variables

#### Is it used? Not often...

- Only relative "degree of belief"
  - > Requires at least two hypotheses
- "Prior" avoidance
- Challenges where single points in parameter space are important
  - > Is sin2b = 0?

## **Some Recommendations**

#### **Define strategy in advance of data analysis**

- Otherwise, significance estimates could and will be biased
- "Blind" analyses can play a role
  - > However, this should not limit the ability to "explore" the data

# Take consistent approach to CL setting & signal measurement

Avoid "flip-flopping" -- F-C offers one approach to this problem

### Describe clearly how you are determining "significance"

#### – Things to remember:

- > Definition of probability
- > Definition of critical region
- > What decisions were taken *a posteriori*?

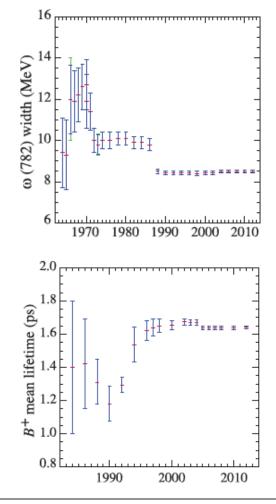
### 4. Blind Analyses

#### To make inferences, we have to assume:

- Random events free from correlation
- More data results in greater precision
- Procedures used are free of bias

#### Are these reasonable assumptions?

- PDG has a set of "history" plots
  - Reveal that some measurements are just wrong
  - Post mortems have indicated that some bias had crept into analysis
    - > Looking for the right answer?
    - > Selection biased by data itself?



### **Piled Higher and Deeper**

#### Piled Higher and Deeper by Jorge Cham

www.phdcomics.com

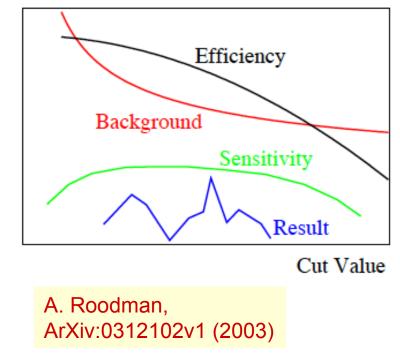


title: "Check it" - originally published 3/31/2014

## **How Can This Happen?**

#### Simple carton illustrates a typical situation

- One is "exploring" the data
- Finds a "cut" that miraculously reduces the background with high efficiency
- But what is the right value of the cut?
- In some cases, it is not so clear
  - Experimenter can make an arbitrary choice
  - But behavioural psychologists claim there is no such thing!



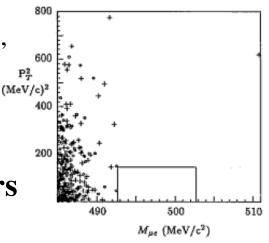
### **Avoiding Experimenter's Bias**

### A standard solution has been to formally "blind" the analysis

- Define *a priori* "signal region" or "measurable" that will not be looked at during analysis
- Define a procedure for "opening the box"

#### **Now been used in HEP for about 20 years**

- Popularized by the BaBar collaboration
  - > committed to using "blind techniques"
- Goes back to 1662 by John Baptista von Helmont
  - > Adopted in the biomedical community as the "gold standard" – double-blind studies – as far back as 1948



### **Too Good to be True?**

### Actually, works pretty well in practice

- Generally accepted as one strategy for reducing the bias

### Some pitfalls/challenges:

- "Blinding" obscures an unanticipated instrumental or theoretical problem
  - > Discover that half the data was missed (true example)!
- After "opening the box", procedure changes because of ancillary studies or measurements
  - > Current example in ATLAS is where
    - Box opened and 5 signal events
    - New "jet cleaning tool being implemented" kills 1 event
    - 17% of background events also reduced, though 9 events in "sideband" all survived
    - Do you use the new "jet cleaning tool"?